



Algorithm Theoretical Basis Document GSICS GEO-AIRS (Details)

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Algorithm Theoretical Basis Document GSICS GEO-AIRS Inter-Calibration



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ATBD Design – Summary



❖GSICS goals require that single pixel collocations anywhere within the GEO field of regard be collected continuously over long term for all bands.

❖GSICS should collect all it can to allow future selection and manipulation by users.



Implementation – Subsetting



```
for each LEO granule
   for each GEO image
                    CLOSE TO GEO NADIR
          if LEO granule is outside of GEO FOR then next granule
                    WITHIN GEO IMAGE
                                                                         Image
          if LEO granule is outside of GEO image then next image
                    CONCURRENT IN TIME
                                                                    Close to GEO Nadar
          time diff = LEO time - GEO time
          if time diff > max sec then
               this image is too early – next image
          else if time diff < -max sec then
                                                                                    Full Disk
               this image (and the rest) is too late – next granule
          else
                                                                    GEO Time
               this image matches the granule – SUBSET
          end
   end (next GEO image)
end (next LEO granule)
```

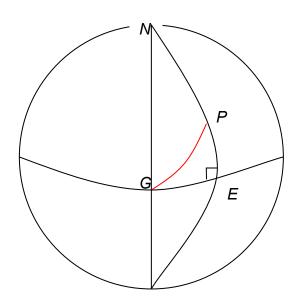
LEO Time & Window

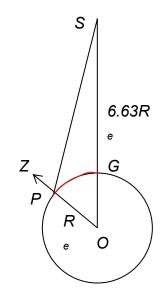


Implementation – Subsetting



Cosine theorem of spherical trigonometry





Cosine and sine theorems of plane trigonometry

$$GP = \angle SOZ$$

$$cos(GP) = cos(gra_ctr_lat)cos(geo_nad_lon - gra_ctr_lon)$$

$$SP^2 = SO^2 + OP^2 - 2 * SO * OP \cos(PG)$$

$$\sin(SPZ) = \sin(PG) * SO / SP$$



Implementation – Collocation



for each LEO pixel

COLLOCATED IN SPACE

if LEO pixel is outside of GEO image then next LEO pixel

CONCURRENT IN TIME

if |LEO_time - GEO_time | > max_sec (300 sec) then next LEO pixel

ALIGNED IN LINE-OF-SIGHT

if $|\cos(\text{geo}_{\text{zen}})/\cos(\text{leo}_{\text{zen}})-1| > \max_{\text{zen}} (0.01)$ then next LEO pixel

UNIFORM ENVIRONMENT

if env stdv > max stdv (10 count) then next LEO pixel

NORMAL GEO FOV

if | fov_mean - env_mean | > (env_stdv/n)(N-n)/(N-1)*G (3) then next LEO pixel

SPATIAL AVERAGING

simple average of GEO pixels in area comparable of LEO FOV

SPECTRAL CONVOLUTION

several choices

OUTPUT THE RESULTS

list of parameters provided

end (next LEO pixel)

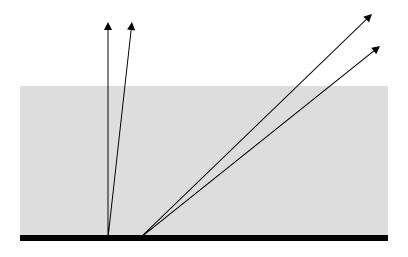


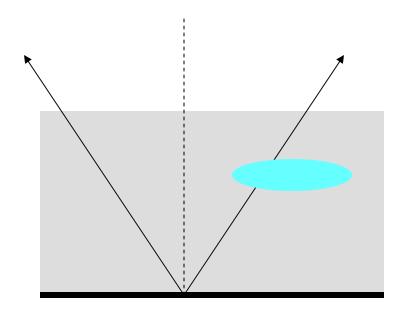
Implementation – Collocation Angle



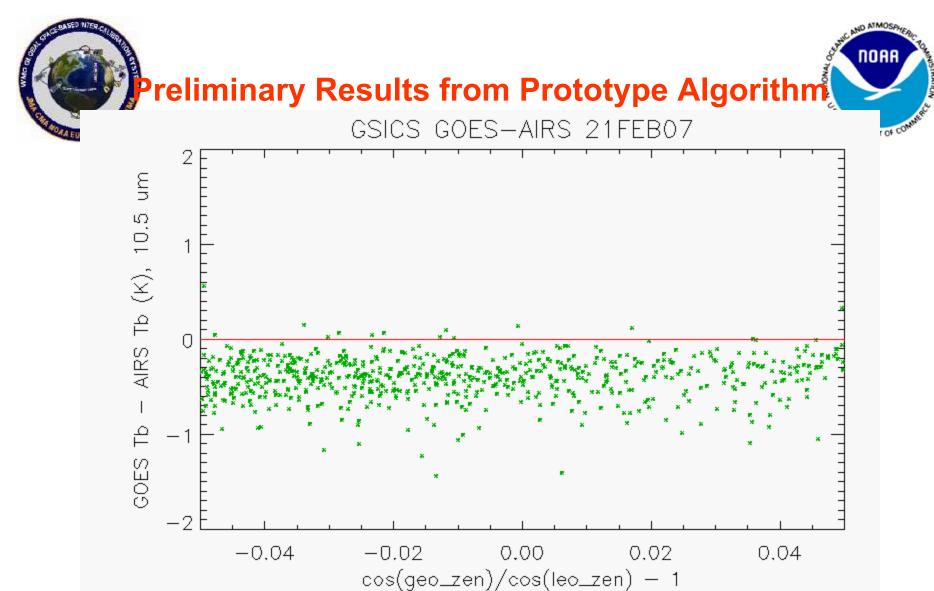
| sec(geo_zen) - sec(leo_zen) |

cos(geo_zen)/cos(leo_zen)-1

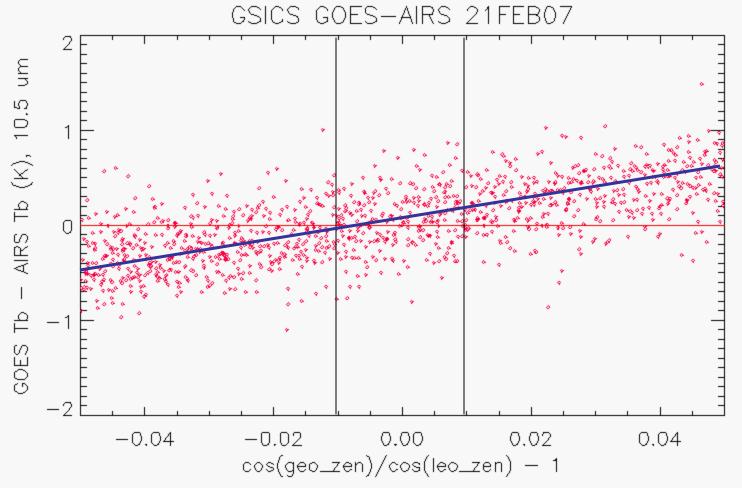




IR channels are often insensitive to difference in azimuth angle



Preliminary Results from Prototype Algorithm



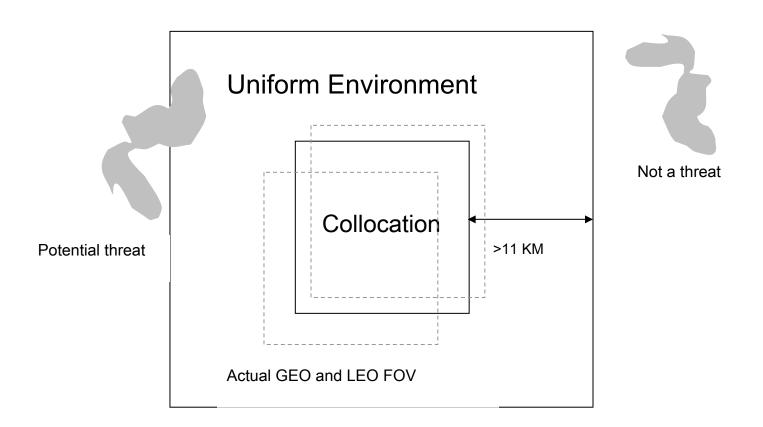
Empirical correction is helpful, although one cannot depend on that too much since this correction depends on the lapse rate

DOAR



Implementation – Collocation Uniform

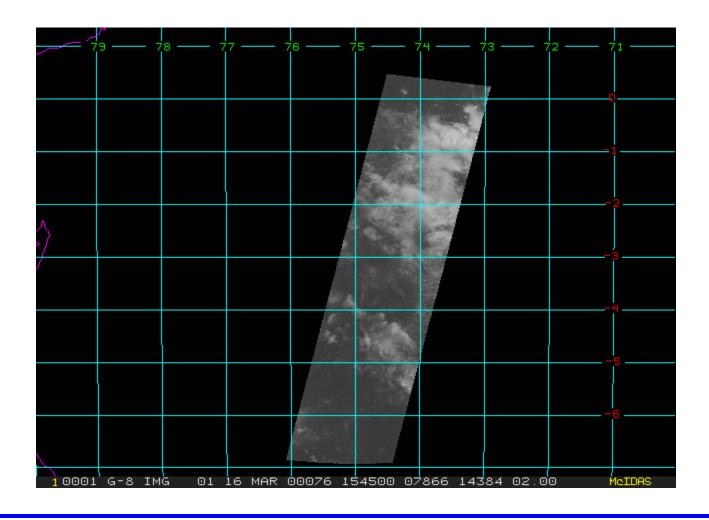






Preliminary Results from Prototype Algorithm

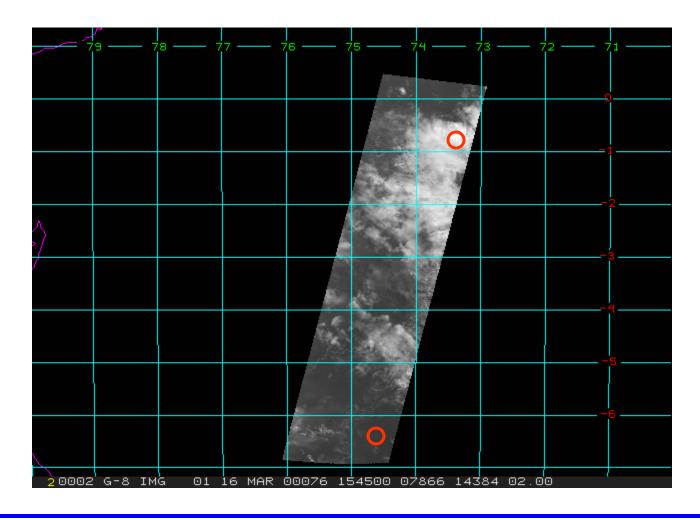






Preliminary Results from Prototype Algorithm







Implementation – Collocation Uniform



- Uniformity measured by standard deviation σ
- $\bullet \sigma$ of T_b is not a good choice.
- $\bullet \sigma$ of radiance is better.
 - Varies with scene T_b and wavelength weighted by mean

Table 1: δT_h in response to 5% δR for GOES IR channels at selected scene temperature.

	3.9 μ m	6.6 μ m	10.7 μ m	12.0 μ m	13.3 μ m
290°K	1.2K	2.0K	3.1K	3.6K	3.9K
250°K	0.9K	1.5K	2.4K	2.7K	2.8K
210°K	0.6K	1.0K	1.5K	1.8K	2.0K





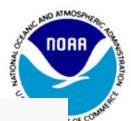


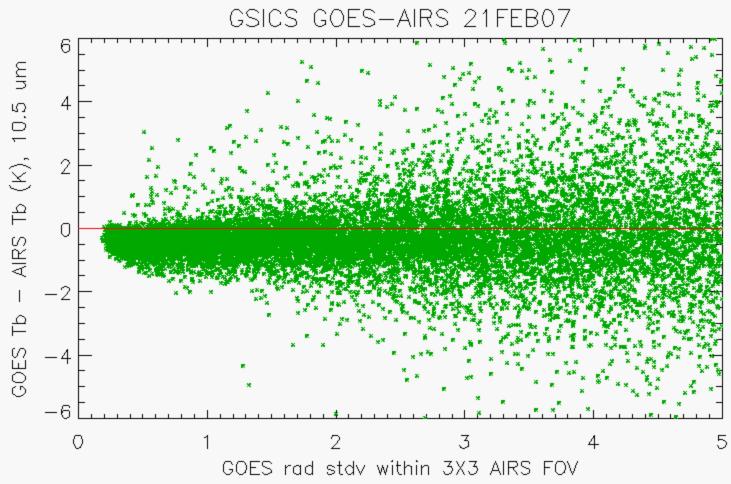
 $\bullet \sigma$ of count is the best

Table 2: δT_{-b} in response to 10 counts for GOES IR channels at selected scene temperature.

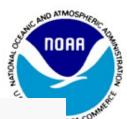
	3.9 μ m	6.6 μ m	10.7 μ m	12.0 μ m	13.3 μ m
290°K	1.6K	0.5K	1.2K	1.2K	1.1K
250°K	7.5K	1.1K	1.9K	1.8K	1.5K
210°K	27K	3.8K	3.7K	3.1K	2.4K

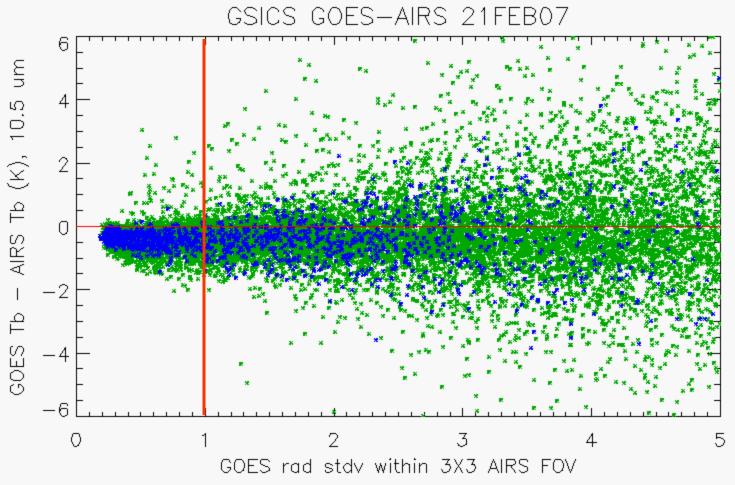








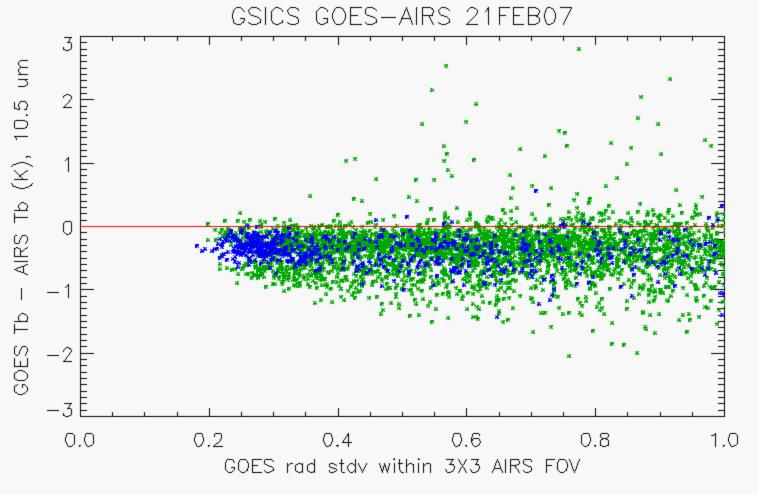




Blue: time difference < 60 seconds





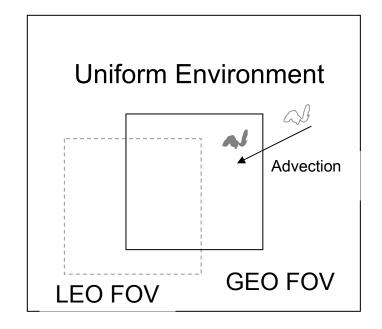








$$\left| \frac{1}{n^2} \sum_{i=1}^{n^2} R_i - M \right| \le \frac{S}{n} \frac{N-n}{N-1} Gaussian (=3)$$





Implementation – Collocation Averaging and Convolution



Spatial average of GEO radiances

$$R_{GEO} = \frac{1}{n^2} \sum_{i=1}^{n^2} R_i$$

Spectral convolution of LEO radiances

$$R'_{GEO} = \frac{\int_{\nu} R_{\nu} \Phi_{\nu} d\nu}{\int_{\nu} \Phi_{\nu} d\nu}$$



Implementation – Output Results



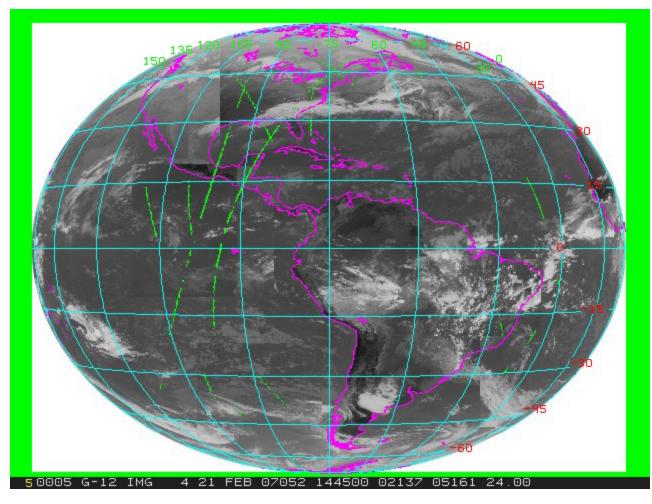
Real * 4	yyddd	year and day of year
Real * 4	hhmmss	hour/minute/sec of GEO observation
Real * 4	time_diff	LEO_time – GEO_time (sec)
Real * 4	zeni_diff	$\cos(\theta \text{GEO})/\cos(\theta \text{LEO})$ - 1
Real * 8	time	LEO time of observation (TAI second)
Real * 4	latitude	collocation latitude (degree east positive)
Real * 4	longitude	collocation longitude (degree north positive)
Real * 4	geo_zen	GEO zenith angle (degree)
Real * 4	leo_zen	LEO zenith angle (degree)
Real * 4	sol_zen	SUN zenith angle (degree)
Real * 4	geo_azi	GEO azimuth angle (degree)
Real * 4	leo_azi	LEO azimuth angle (degree)
Real * 4	sol_azi	SUN azimuth angle (degree)
Real * 4	airs_cnv_shift	Ch6 shift SRF (irrelevant in general – to be deleted)
Real * 4	airs_mmg_shift	Ch6 shift SRF (irrelevant in general – to be deleted)
Real * 4	stat(6,4)	mean & stdv of collocation environment, mean & stdv of
		collocation target, convoluted AIRS radiance using modified
Real * 4	loo_rod(2278)	Kato and Gunshor methods, for four channels
	leo_rad(2378)	AIRS spectral radiances at 2378 channels
Real * 4	geo_rad(17,9,4)	GEO rad at 17 elements, 9 lines, and 4 channels

Further discussion after break





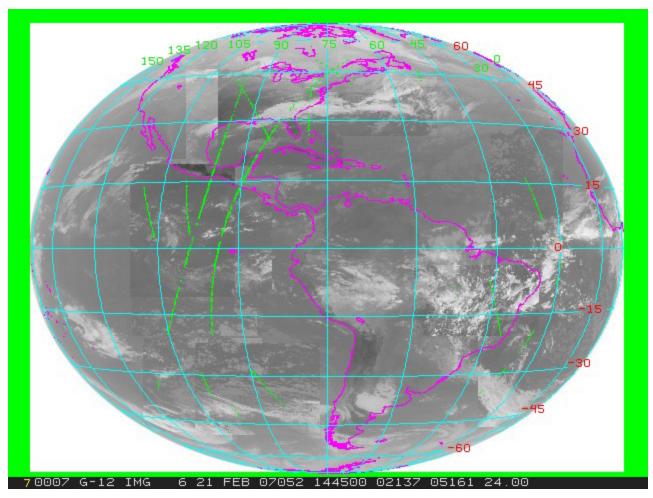












All bands data saved even if only one band qualifies for collocation